

Universal Scaling Laws in Quantum Quench

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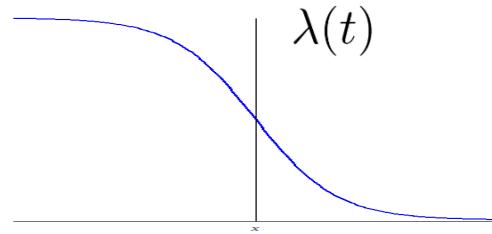
(*Pawel Caputa, S.R.D. and Masahiro Nozaki*)

(*S.R.D., S. Hampton and Si-Nong Liu*)

(*P. Caputa, D. Das, S.R.D*)

Excited States and Quantum Quench

- An efficient way to produce excited states in quantum many-body systems is to subject it to a **quantum quench**.
- Typically one can start with the ground state of an initial (*time independent*) Hamiltonian, and then introduce **time dependence in a coupling** so that at late times one gets another time-independent Hamiltonian.



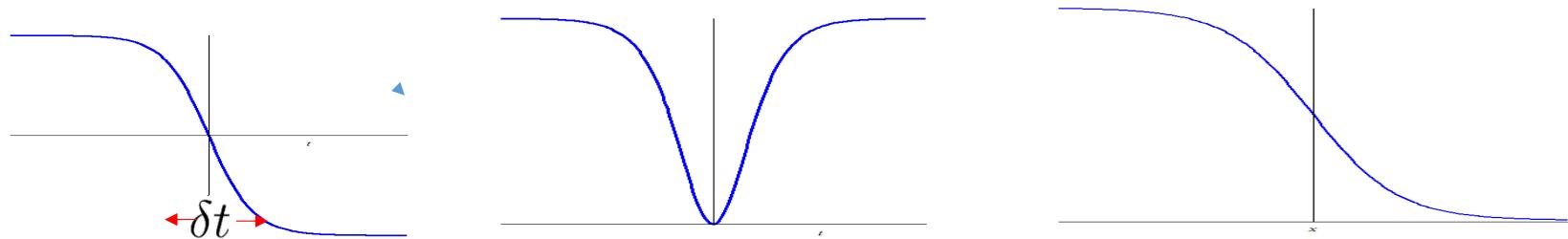
- The (*Heisenberg picture*) state is then an excited state of the new Hamiltonian.

- This excited state then evolves non-trivially in time and we are typically interested in a variety of questions about the resulting non-equilibrium state.
- One set of questions relate to **thermalization** : how does the system reach a steady state, and in what sense is this state (approximately) **thermal** ?
- Another set of questions relate to **universal properties** of various quantities for situations where the quench involves the vicinity of critical points. These questions sometimes involve early time behavior – in the critical region.
 - An important aspect of this is the **scaling of expectation values** with the quench rate, the exponents being determined by the critical point.
- In this talk I will explore some **well known** as well as some **newly discovered** scaling properties.

Scaling in Critical Quantum Quench

- Quantum quench involving critical points can be described by an action

$$S = S_{critical} - \int dt \int d^{d-1}x \lambda(t) \mathcal{O}(\vec{x}, t) \quad \lambda(t) = \lambda_0 F(t/\delta t)$$



- Near the critical point the coupling has the behavior

$$\lambda(t) = \lambda_0 (t/\delta t)^r$$

Kibble Zurek Scaling

- When the quench is **slow** compared to the physical mass scales in the problem

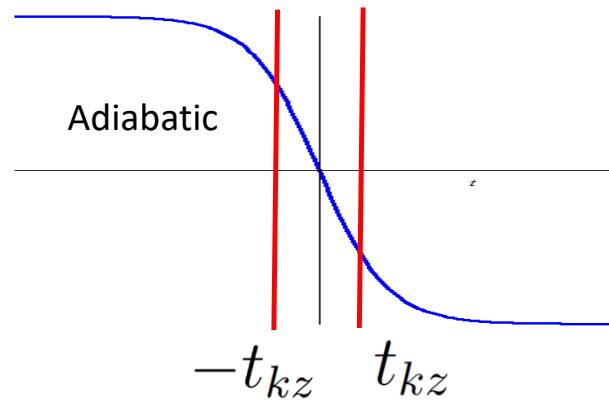
$$\delta t \gg \lambda_0^{-1/(d-\Delta)}$$

- Kibble and Zurek proposed that after the time $-t_{kz}$ when **adiabatic evolution breaks down** the system becomes, to lowest approximation, **diabatic**.
- The instantaneous correlation length at $-t_{kz}$ is the **only length scale** in the critical region

$$t_{kz} \sim \left(\frac{\delta t}{\lambda_0^{1/r}} \right)^{\frac{z\nu}{z\nu+1}}$$

ν = Correlation length exponent

z = Dynamical critical exponent



$$\xi_{kz} = \xi(t_{kz}) = \left(\frac{\delta t}{\lambda_0^{1/r}} \right)^{\frac{\nu}{z\nu+1}}$$

- The **instantaneous energy gap** is given by

$$E_{gap}(t) \sim [\lambda_0(\frac{t}{\delta t})^r]^{z\nu}$$

- The **time at which adiabaticity fails** is given by Landau criterion

$$[\frac{1}{E_{gap}(t)^2} \frac{dE_{gap}(t)}{dt}]_{t=t_{KZ}} \sim 1$$

- This immediately leads to

$$t_{KZ} \sim (\frac{\delta t}{\lambda_0^{1/r}})^{\frac{z\nu}{z\nu+1}}$$

- If the system is **frozen in the time interval between $-t_{KZ}$ and t_{KZ}** , to lowest order the expectation value of the operator during these times would be the same as that at time t_{KZ}
- However the latter is still given by the **adiabatic approximation**. This in turn has to scale according to its operator dimension

$$\langle \mathcal{O} \rangle \sim [\xi(-t_{KZ})]^{-\Delta} \sim (\delta t)^{-\frac{\Delta\nu}{z\nu+1}}$$

- An improved version involves **scaling functions** for correlators

$$\langle \mathcal{O}(t) \rangle \sim \xi_{kz}^{-\Delta} f(t/t_{kz})$$

$$\langle \mathcal{O}(\vec{x}, t) \mathcal{O}(\vec{x}', t') \rangle \sim \xi_{kz}^{-2\Delta} g\left(\frac{|\vec{x} - \vec{x}'|}{\xi_{kz}}, \frac{|t - t'|}{t_{kz}}\right)$$

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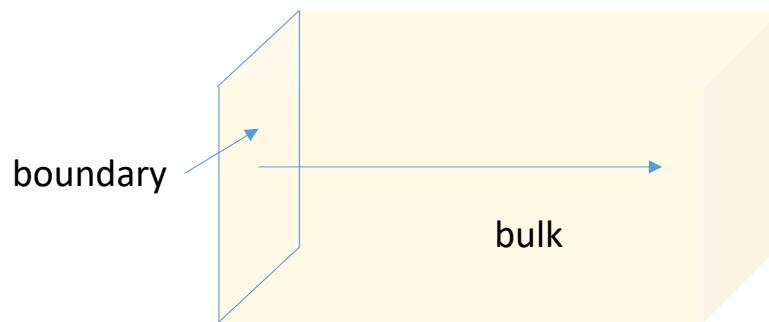
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- The initial assumptions of Kibble-Zurek appear too simplistic – nevertheless **the scaling results hold** for many models.
- Kibble proposed this for **thermal transitions** in the context of **cosmology** – Zurek extended this to **condensed matter systems** - there is some experimental evidence for this. Recently generalized for quantum transitions.
- Experiments in cold atom systems seem to show this kind of scaling.
- However there is little understanding of this in quantum field theory – There is no good understanding of the **renormalization group for time dependent systems**.

The Holographic Setup

- The AdS/CFT correspondence provides a duality between large N quantum field theories which are deformations of conformal field theories, and theories which contain gravity in asymptotically anti-de Sitter spacetimes.
- The QFT lives on the boundary of AdS space-time. The dual gravitational theory lives in the “bulk” – the interior of AdS
- In many situations, when the QFT is strongly coupled, the dual gravitational description is classical General Relativity with other appropriate bulk fields – a strongly coupled problem in QFT then gets mapped to a classical problem.



- In this regime of strong coupling **operators in the field theory** are dual to **fields in the bulk**

scalar $\mathcal{O}(\vec{x}, t) \leftrightarrow \phi(\vec{x}, t, z)$ scalar field

vector current $J_\mu(\vec{x}, t) \leftrightarrow A_\mu(\vec{x}, t, z)$ gauge field

EM tensor $T_{\mu\nu}(\vec{x}, t, z) \leftrightarrow h_{\mu\nu}(\vec{x}, t, z)$ metric perturbation

- Near the boundary one has an asymptotic expansion

$$\phi(\vec{x}, t, z) \sim z^{d-\Delta} [\lambda(\vec{x}, t) + O(z^2)] + z^\Delta [A(\vec{x}, t) + O(z^2)]$$

- The integration function $\lambda(\vec{x}, t)$ is **the coupling** for the deformation

$$S = S_{CFT} - \int d^{d-1}x dt \lambda(\vec{x}, t) \mathcal{O}(\vec{x}, t)$$

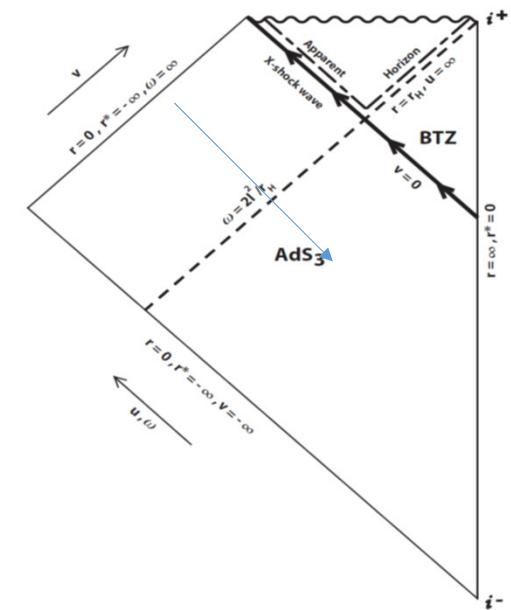
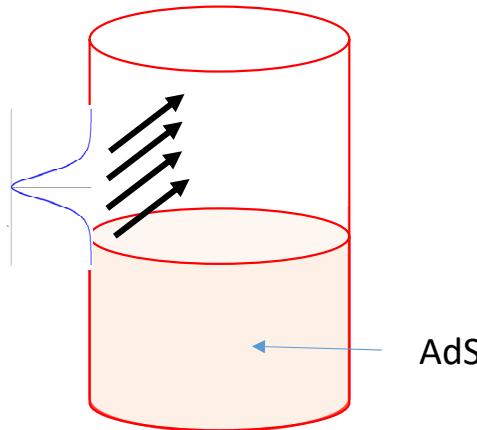
- While the integration function $A(\vec{x}, t)$ gives **the response**

$$\langle \mathcal{O}(\vec{x}, t) \rangle = A(\vec{x}, t)$$

- The problem of **quantum quench** translates into a classical bulk problem with a **time dependent boundary condition**.
- An important aspect of this relates to thermalization of the field theory.

This becomes dual to **formation of black holes** in the bulk.

(*Danielsson, Keski-Vakkuri and Kruczenski; Chesler and Yaffe; Bhattacharyya and Minwalla; Balasubramanian et.al; Bizon & Rostworoski; Buchel, Lehners, Leibling; Basu and Krishnan; Craps, Evnin, van Hoof;.....*)



- Here we will study quantum quench involving a critical point – concentrating on universal scaling behavior
 - (1) Identify a bulk theory which is dual to a critical point.
 - (2) Impose time dependent boundary conditions on the bulk field which is dual to a relevant operator which crosses or approaches the critical point.
 - (3) Calculate the response

Holographic Kibble-Zurek

This has been studied both at zero and non-zero temperature

- *P. Basu and S.R.D., JHEP 1201 (2012) 103*
P. Basu, D. Das, S.R.D. & T. Nishioka, JHEP 1301 (2013) 107
P. Basu, D. Das, S.R.D. & K. Sengupta, JHEP 1311 (2013) 186
S.R.D. & T. Morita, JHEP 1501 (2015) 084.
- *J. Sonner, A. del Campo and W. Zurek, Nat. Comm. 6 (2015) 7405*
- *P. Chesler, A. Garcia-Garcia and H. Liu, Phys. Rev. X (2015) 021015*

- There are many **holographic realizations** of quantum critical phase transitions which describe superfluidity, magnetism etc.
- In these models, the transition is obtained by tuning some **coupling of the boundary field theory**, e.g. a **chemical potential**. In this case the bulk meaning of this chemical potential would be the **boundary value of a bulk gauge field**.

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- In most cases, the critical point corresponds to the **appearance of a zero mode** of the corresponding bulk field at the linearized level.
- When the chemical potential is larger than this critical value, the **trivial solution becomes unstable**, and the **full nonlinear equations have a stable solution** which is non-trivial – this is manifested as a **nonzero expectation value of the corresponding operator**.

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- When the chemical potential is larger than this critical value, the **trivial solution becomes unstable**, and the **full nonlinear equations have a stable solution** which is non-trivial – this is manifested as a **nonzero expectation value of the corresponding operator**.
- To study **quantum quench** we make this **boundary condition time dependent** and solve the **bulk equations of motion** – from this read off the response in the field theory.
- We want to choose the time dependence such that the **coupling crosses a critical point**.

- We implemented this in several examples of quench across **holographic quantum critical points**.
- In all these cases, we showed analytically that in the slow quench regime this **zero mode dominates the dynamics in the critical region**.
- The **zero mode dynamics has a scaling solution** – this leads to **Kibble-Zurek scaling**.

Holographic Fast Quench

- *Buchel, Lehner, Myers and van Niekerk* looked at the opposite regime of a **fast but smooth quench**, $\delta t \ll m_{gap}^{-1}$, using these **holographic** setups.
- They found a **different kind of scaling behavior** for quenches driven by bulk scalars

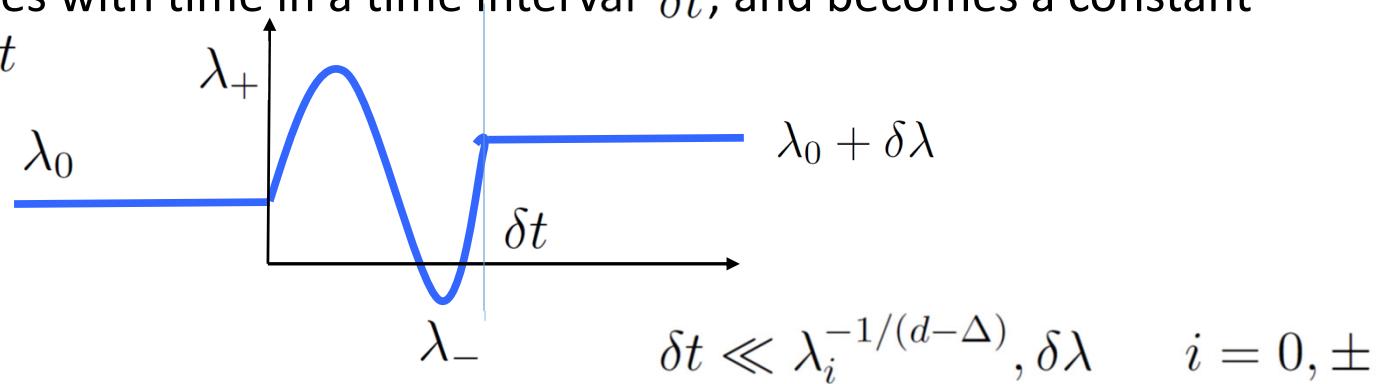
$$\langle \mathcal{O} \rangle \sim (\delta t)^{d-2\Delta} \quad \langle \mathcal{E} \rangle \sim (\delta t)^{d-2\Delta} \rightarrow \infty$$

- These results, first obtained **numerically**, can be understood by noticing that in this regime the response is dominated by the space-time close to the boundary.

- It turns out that this is in fact a result in **any relativistic quantum field theory**, regardless of holography.
(S.R.D., Damian Galante and Robert C. Myers)
- Consider **a general interacting CFT deformed by some relevant operator** with a **time dependent coupling**

$$S = S_{CFT} - \int dt \int d^{d-1}x \lambda(t) \mathcal{O}(\vec{x}, t) \quad \lambda(t) = \lambda_0 F(t/\delta t)$$

- The coupling is at some constant value λ_0 for $t < 0$. At $t = 0$ it smoothly turns on and changes with time in a time interval δt , and becomes a constant quickly after $t = \delta t$



- Start computing $\langle \mathcal{O} \rangle$. in **perturbation theory**. The first few terms are

$$\begin{aligned}\langle \mathcal{O}(\vec{x}, t) \rangle - \langle \mathcal{O}(\vec{x}, t) \rangle_{\lambda_0} &= -\delta\lambda \int_0^t dt' F(t'/\delta t) \int d^{d-1} \vec{x}' G_{R, \lambda_0}(\vec{x} - \vec{x}', t - t') \\ &\quad + \frac{\delta\lambda^2}{2} \int_0^t dt' F(t'/\delta t) \int d^{d-1} \vec{x}' \int_0^t dt'' F(t''/\delta t) \int d^{d-1} \vec{x}'' K_{\lambda_0}(t', \vec{x}'; t'', \vec{x}''; t, \vec{x})\end{aligned}$$

- Where $G_{R, \lambda_0}(\vec{x}, t)$ is the **retarded Green's function of the initial theory**

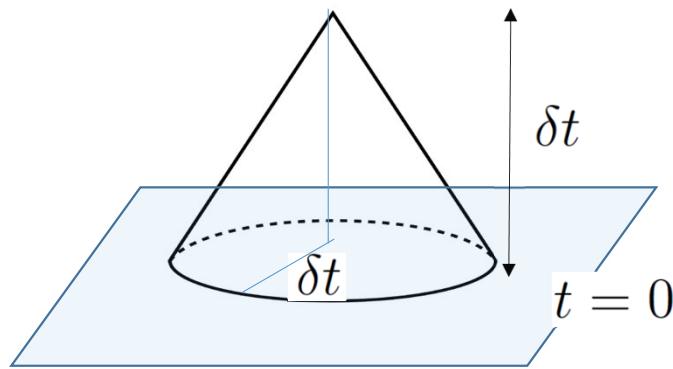
$$G_{R, \lambda_0}(\vec{x}, t) = i\theta(t) \langle 0 | [\mathcal{O}(\vec{x}, t), \mathcal{O}(0, 0)] | 0 \rangle_{\lambda_0}$$

- $K_{\lambda_0}(t', \vec{x}'; t'', \vec{x}''; t, \vec{x})$ is a three point function.
- All quantities above are **renormalized** quantities. This requires counter-terms which depend on $\lambda(t)$ as well as its time derivatives : these counter-terms can be read off, with precise coefficients, from **adiabatic expansions**

- Consider the first term

$$\int_0^t dt' F(t'/\delta t) \int d^{d-1} \vec{x}' G_{R,\lambda_0}(\vec{x} - \vec{x}', t - t')$$

- While the integration over \vec{x}' has been written as over entire space, **causality** implies that **only the region** $|\vec{x} - \vec{x}'| \leq t$ **has a non-trivial contribution**.
- Now suppose we want to calculate the quantity at $t = \delta t$, **right at the end of the quench**. Then both the space and time intervals which appear in the integral are at most of size δt



- Recall that the scale δt is smaller than all other physical scales in the problem, in particular the scale associated with the deformation of the CFT by λ_0
- Therefore the Green's function $G_{R,\lambda_0}(\vec{x} - \vec{x}', t - t')$ is basically the Green's function of the UV **conformal field theory**.

$$G_{R,\lambda_0}(\vec{x}' - \vec{x}, t' - t) \sim G_{R,CFT}(\vec{x}' - \vec{x}, t' - t) \quad |\vec{x}' - \vec{x}|, |t' - t| \ll (\lambda_0)^{-1/(d-\Delta)}$$

- This means that in the leading contribution, the only scale which appears in the integral is δt

- This leads to an expression for the response which is of the form

$$\langle \mathcal{O}_\Delta(t) \rangle_{\text{ren}} - \langle \mathcal{O}_\Delta(t) \rangle_{\text{ren},\lambda_0} = (\delta t)^{-\Delta} [b_1(t/\delta t) g + b_2(t/\delta t) g^2 + \dots]$$

- Where we have introduced a **dimensionless coupling** $g \equiv \delta \lambda \delta t^{d-\Delta}$
- In the **fast quench limit** this dimensionless coupling is small, so that to leading order we get the universal scaling

$$\langle \mathcal{O} \rangle \sim (\delta t)^{d-2\Delta}$$

- There is a similar argument for the **energy density** produced. Since the coupling remains constant after δt this quantity is in fact the net energy produced.

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- There is a similar argument for the **energy density** produced. Since the coupling remains constant after δt this quantity is in fact the net energy produced.
- Note that the result is completely general and **depends only on the properties of the conformal field theory in the UV**.
- A more detailed understanding of this general result appears in [*Berentsein and Miller*](#) and by [*Dymarsky and Smolkin*](#).
- Explicit calculations in solvable (free) field models.

The Abrupt Limit

- It may appear puzzling that in the $\delta t \rightarrow 0$ limit, the fast quench formulae like

$$\langle \mathcal{O} \rangle \sim (\delta t)^{d-2\Delta}$$

diverge for $2\Delta > d$. This should be the abrupt quench limit !

- In this limit several results are known in low dimensions – mostly due to *Calabrese and Cardy*. The results are certainly not infinite.
- The reason of course is that we have been considering **renormalized quantities**, so the quench rate is **fast compared to physical mass scales**, but **slow compared to the UV scale**.

$$\Lambda_{UV}^{-1} \ll \delta t \ll m_{gap}^{-1}$$

- To understand this better we have studied solvable models on a lattice.

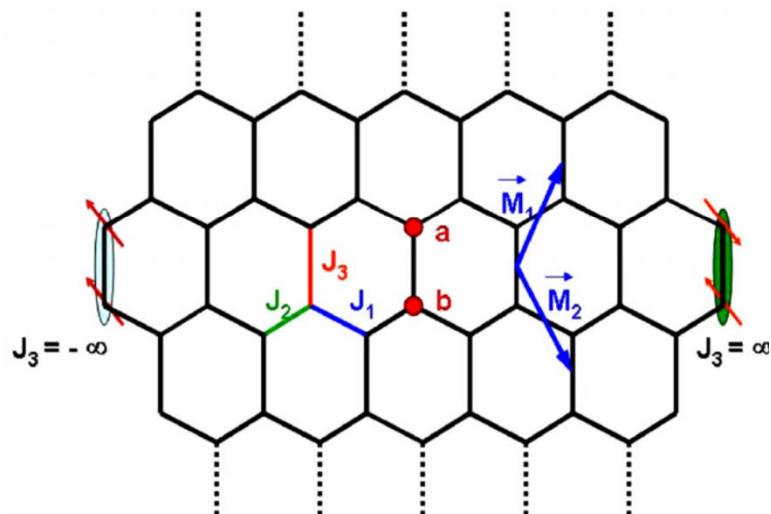
Models on a lattice

(D.Das, S.R.D., D. Galante, R. Myers and K. Sengupta)

- We have found exactly solvable quench protocols (with couplings which asymptote to constant values) for several solvable lattice models.



$$H = - \sum_n [g(t)\sigma^3(n) + \sigma^1(n)\sigma^1(n+1)] \quad \text{ISING}$$



$$H_{\text{Kitaev}} = \sum_{j+l=\text{even}} \left[J_1 \sigma_{j,l}^{(1)} \sigma_{j+1,l}^{(1)} + J_2 \sigma_{j,l}^{(2)} \sigma_{j-1,l}^{(2)} + J_3 \sigma_{j,l}^{(3)} \sigma_{j,l+1}^{(3)} \right]$$

KITAEV HONEYCOMB

- Both the models can be written in terms of **fermions** using Jordan Wigner

$$H = \int \frac{d^D k}{(2\pi)^D} \chi^\dagger(\vec{k}) \left[-m(\vec{k}, t) \sigma_3 + G(\vec{k}) \sigma_1 \right] \chi(\vec{k}) \quad \chi(\vec{k}) = \begin{pmatrix} \chi_1(\vec{k}) \\ \chi_2(\vec{k}) \end{pmatrix}$$

	d	$m(k, t)$	$G(k)$	
Ising	2	$g(t) - \cos k$	$\sin k$	$\chi_2(\vec{k}) = \chi_1^\dagger(-\vec{k})$
Kitaev	3	$-2(J_3(t) + J_1 \cos k_1 + J_2 \cos k_2)$	$2(J_1 \sin k_1 - J_2 \sin k_2)$	

- The **Ising Model** has an **isolated critical point**. The RG flow from this UV fixed point gives a continuum theory of a **relativistic Majorana fermion** in **1+1 dimensions**

$$g = 1 - \epsilon(t)$$

- Introduce a lattice spacing, rescale momenta, couplings

$$p = k/a, \quad m(t) = \epsilon(t)/a, \quad \text{and} \quad \psi(p) = a^{1/2} \hat{\chi}(a p) \quad \hat{J} = J a$$

- Leads to the QFT Hamiltonian

$$H_{\text{Ising}}^{\text{cont}} = 2\hat{J} \int_0^\infty \frac{dp}{2\pi} \psi^\dagger(p) [m(t) \sigma_3 + p \sigma_1] \psi(p)$$

- Thus a time dependent transverse field of the Ising model is essentially a **time dependent mass** for a Majorana fermion.

Exactly Solvable Quench Protocol

- Even though the Hamiltonians look quite complicated, it turns out that we can find **exact analytic solutions** of the operator equations of motion for interesting profiles

$$\begin{pmatrix} g(t) \\ J_3(t) \end{pmatrix} = a + b \tanh(t/\delta t)$$

- Recall the Hamiltonian is of the form

$$H = \int \frac{d^D k}{(2\pi)^D} \chi^\dagger(\vec{k}) \left[-m(\vec{k}, t) \sigma_3 + G(\vec{k}) \sigma_1 \right] \chi(\vec{k})$$

- The **positive and negative frequency** “in” solutions are

$$U(\vec{k}, t) = \begin{pmatrix} -i\partial_t + m(\vec{k}, t) \\ -G(\vec{k}) \end{pmatrix} \phi(\vec{k}, t)$$

$$V(\vec{k}, t) = \begin{pmatrix} G(\vec{k}) \\ i\partial_t + m(\vec{k}, t) \end{pmatrix} \phi^*(\vec{k}, t)$$

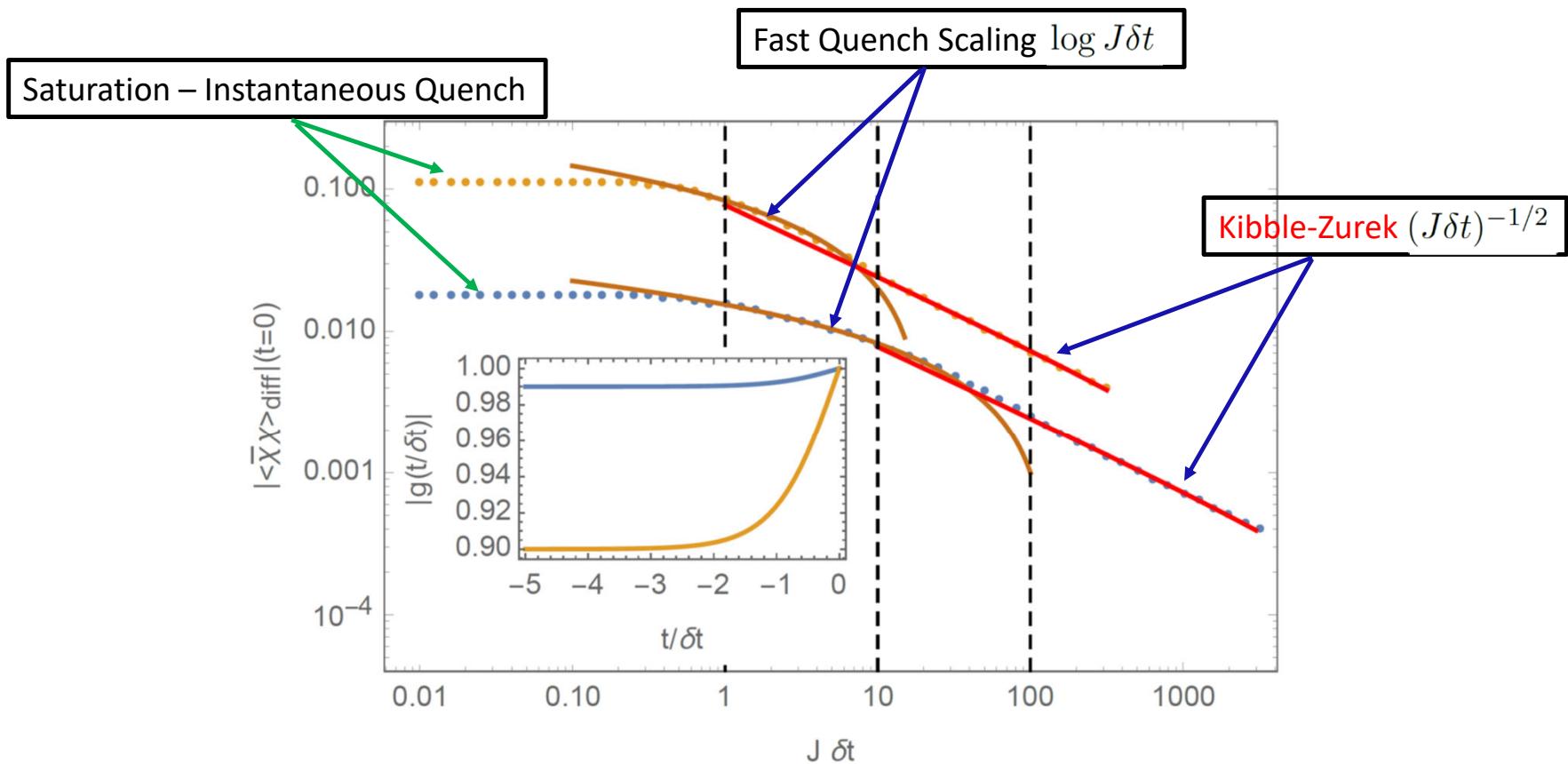
- This is like solving the [Dirac equation in momentum space](#) with a “mass” which is both **momentum dependent** and **time dependent**

$$m(k, t) = A(\vec{k}) + B \tanh(t/\delta t)$$

	$A(\vec{k})$	B
Ising	$-2J(\cos k - a)$	$2Jb$
Kitaev	$-2J(\cos k_1 + \cos k_2 - 2a)$	$4Jb$

- The mode functions can be solved exactly for all δt in terms of hypergeometric functions.
- So we can express the [response](#) ${}_{in}\langle 0|\bar{\chi}\chi|0\rangle_{in}$ as a [momentum integral](#) which we evaluate numerically

Ising Model : Small Amplitude



- The **saturation** happens when **all the modes in the Brillouin zone become non-adiabatic**. For small amplitude this happens when the **quench rate is at the scale of the lattice spacing**, $J\delta t \lesssim 1$ **independent of the amplitude**.

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- For **rates slow compared to physical scales** one gets, as expected, a **Kibble Zurek scaling** regime with exponents predicted by the continuum theory

Scaling of non-local quantities

- It turns out that several interesting **non-local** quantities which characterize quantum systems also show these three regimes of scaling.

Entanglement Entropy (P. Caputa, S.R.D., M. Nozaki, A. Tomiya)

Complexity (P. Caputa et.al.; S. Liu)

- It appears that this universality is shared by many quantities.

Non-relativistic theories

- Very often experimental situations involve **non-relativistic theories** –e.g. atoms in optical traps.
- We would like to learn what happens to these universal scalings in such non-relativistic situations.
- In particular, the argument for **fast quench** scaling heavily relied on **relativistic invariance** – one might wonder if there is some version of this in non-relativistic systems as well.
- Many such systems (typically those on a lattice) have an effective light cone provided by the **Lieb Robinson bound** – so there is a possibility that we may get such scaling regimes as well.
- We will try to work our way up by finding **exactly solvable models of quench** – hopefully this will teach us some lessons as it did for relativistic theories.

Non-relativistic Fermions in $\pm x^2$ potentials

- We will consider a system of **N** non-relativistic fermions in **1+1 dimensions** with the Hamiltonian

$$H = \int dx \psi^\dagger(x) \left[-\frac{\hbar^2}{2m(t)} \frac{\partial^2}{\partial x^2} \pm \frac{1}{2} m(t) \nu^2(t) x^2 \right] \psi(x)$$

- The + sign is relevant for **harmonic traps**. We will see later that the – sign will be relevant for **two dimensional strings**.
- We have the standard anti-commutators and the **particle number constraint**

$$\{\psi(x), \psi^\dagger(x')\} = \delta(x - x') \quad \int_{-\infty}^{\infty} \psi^\dagger(x, t) \psi(x, t) = N$$

- Heisenberg equations of motion is the Schrodinger equation with appropriate initial conditions

$$i \frac{\partial \psi(x, t)}{\partial t} = H \psi(x, t)$$

- A general solution of the Schrodinger problem can be mapped to a solution of a non-linear equation as follows.
- First introduce a **new time variable** and a **scaled frequency**

$$d\tau = \frac{dt}{m(t)} \quad \omega(\tau) = m(t)\nu(t)$$

- This maps the problem to **a constant mass** and a **time dependent frequency**

$$i\frac{\partial\psi(x,\tau)}{\partial\tau} = [-\frac{1}{2}\frac{\partial^2}{\partial x^2} \pm \frac{1}{2}\omega^2(t)x^2]\psi(x,\tau)$$

- This problem has been studied (*Lewis and Risenfeld*; *Pedrosa*; *Ji, Kim and Kim*). Make a change of variables

$$\begin{aligned}\tau &\rightarrow T = \int^{\tau} \frac{d\tau'}{\rho(\tau')^2} \\ x &\rightarrow y = \frac{x}{\rho(\tau)}\end{aligned}$$

- Define a new field $\Phi(y, T)$ defined by

$$\psi(x, \tau) = \frac{1}{\sqrt{\rho(\tau)}} \exp\left[-\frac{i}{2} \frac{\partial_\tau \rho(\tau)}{\rho(\tau)}\right] \Phi(y, T)$$

- Obeys the Schrodinger equation

$$i \frac{\partial \Phi(y, T)}{\partial T} = \left[-\frac{1}{2} \frac{\partial^2}{\partial y^2} \pm \frac{1}{2} y^2 \right] \Phi(y, T)$$

- If the function $\rho(\tau)$ obeys a generalized Ermakov-Pinney equation

$$\partial_\tau^2 \rho(\tau) \pm \omega(\tau)^2 \rho(\tau) = \pm \frac{1}{\rho(\tau)^3}$$

- So the problem of solving the Schrodinger equation with a time dependent mass and a time dependent frequency can be reduced to the equation with constant mass and frequency – and a solution of this nonlinear equation.

- A general solution of the Ermakov-Pinney equation is of the form

$$\rho(\tau)^2 = Af(\tau)^2 + 2Bf(\tau)g(\tau) + Cg(\tau)^2$$

- Where $f(\tau), g(\tau)$ are two linearly independent solutions of

$$\partial_\tau^2 X \pm \omega(\tau)^2 X = 0$$

- And the constants are related by

$$AC - B^2 = \pm \frac{1}{Wr(f, g)^2}$$

- We have therefore reduced the solution of the problem to solutions of the **classical equations of motion for the same Hamiltonian.**

- Using the solution of the field equation one can now quantize the theory canonically

$$\psi(x, t) = \sum_{n=0}^{\infty} a_n \psi_n(x, t)$$

$$\{a_m, a_n^\dagger\} = \delta_{mn} \quad \{a_n, a_m\} = \{a_m^\dagger, a_n^\dagger\} = 0$$

- For quench protocols which asymptote to constant frequencies at early and late times the “in” modes have the property

$$\text{Lim}_{t \rightarrow \infty} \psi_n(x, t) \sim e^{-i\alpha t} \quad \alpha > 0$$

- While the function $\rho(\tau)$ must have the initial condition

$$\text{Lim}_{\tau \rightarrow -\infty} \rho(\tau) = \rho_{in} = \frac{1}{\sqrt{m_{in}\nu_{in}}} = \frac{1}{\sqrt{\omega_{in}}}$$

- This defines the Heisenberg picture “in” vacuum state

$$\begin{aligned} a_n |in\rangle &= 0 & n \geq N \\ a_n^\dagger |in\rangle &= 0 & 0 \leq n \leq N-1 \end{aligned}$$

Observables

- In fact many interesting observables can be expressed entirely in terms of the function $\rho(\tau)$
- Consider for example the expectation value of the quenched operator

$$\mathcal{O} = \int_{-\infty}^{\infty} x^2 \psi^\dagger(x, \tau) \psi(x, \tau)$$

- In the “in” state this is

$$\begin{aligned} \langle 0 | \mathcal{O}(\tau) | 0 \rangle &= \sum_{n=0}^{N-1} \int_{-\infty}^{\infty} dx \ x^2 \ \psi_n^*(x, \tau) \psi_n(x, \tau) \\ &= \rho^2(\tau) \sum_{n=0}^{N-1} \int_{-\infty}^{\infty} dY Y^2 \phi_n^*(Y) \phi_n(Y) = \rho(\tau)^2 \sum_{n=0}^{N-1} (n + 1/2) = \boxed{\frac{N^2}{2} \rho(\tau)^2} \end{aligned}$$

Entanglement Entropy

- Another interesting quantity is the **entanglement entropy** of some region A.
- For free fermions it turns out that this quantity can be expressed in terms of **cumulants of the particle number distribution**. (*Kilch and Levitov*)
- For large N the leading contribution is the dispersion

$$S_A(\tau) = \frac{\pi^2}{3} [\langle N_A(\tau)^2 \rangle - \langle N_A(\tau) \rangle^2]$$

$$N_A(\tau) = \int_A dx \psi^\dagger(x, \tau) \psi(x, \tau)$$



- In the “in” state this can be further simplified (*Collura, Sotiriadis & Calabrese*)

$$S_A(\tau) = \langle in | N_A(\tau) | in \rangle - \int_A dx \int_A dy |C(x, y, \tau)|^2$$

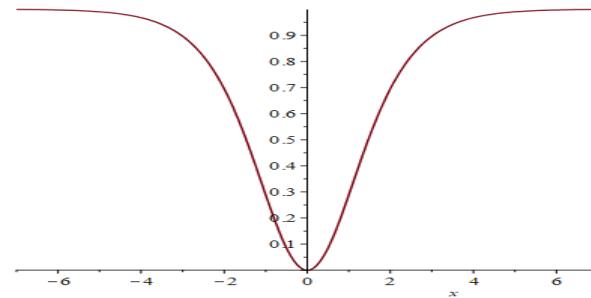
$$C(x, y, \tau) = \langle in | \psi^\dagger(x, \tau) \psi(y, \tau) | in \rangle$$

- Using the mapping described above this can again be expressed in terms of the results for fermions in a potential with constant frequency, but with **an interval rescaled by the function $\rho(\tau)$**

$$S_A[\omega(\tau)] = S_{A_P}[\omega = 1] \quad A_P : \frac{a}{\rho(\tau)} \leq x \leq \frac{b}{\rho(\tau)}$$

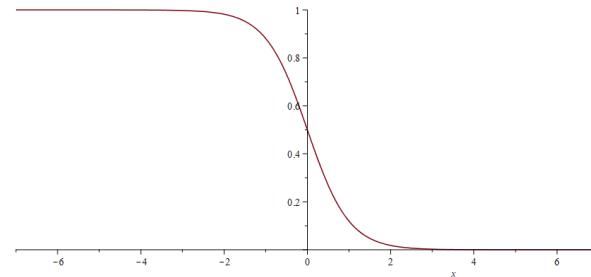
- We have investigated explicit solutions for two interesting protocols
- The first is a “**Cis-Critical Protocol (CCP)**”

$$\omega(\tau)^2 = \omega_0^2 \tanh^2(t/\delta t)$$



- The second is a End-Critical protocol (ECP)

$$\omega^2(\tau) = \frac{1}{2}\omega_0^2 \left(1 - \tanh \frac{\tau}{\delta t}\right)$$



- We have investigated explicit solutions for two interesting protocols
- The first is a “**Cis-Critical Protocol (CCP)**”

$$\omega(\tau)^2 = \omega_0^2 \tanh^2(t/\delta t)$$

- The “in” classical solution is

$$f_{CCP}(\tau) = \frac{1}{\sqrt{2\omega_0}} \frac{2^{i\omega_0\delta t} \cosh^{2\alpha}(\tau/\delta t)}{E_{1/2}\tilde{E}'_{3/2} - E'_{1/2}\tilde{E}_{3/2}} \times \\ \left\{ \tilde{E}'_{3/2} {}_2F_1(a, b; \frac{1}{2}; -\sinh^2 \frac{\tau}{\delta t}) + E'_{1/2} \sinh \frac{\tau}{\delta t} {}_2F_1(a + \frac{1}{2}, b + \frac{1}{2}; \frac{3}{2}; -\sinh^2 \frac{\tau}{\delta t}) \right\}$$

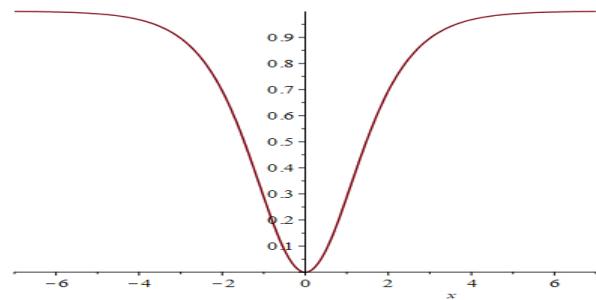
- where

$$\alpha = \frac{1}{4}[1 + \sqrt{1 - 4\omega_0^2\delta t^2}]$$

$$a = \alpha - \frac{i}{2}\omega_0\delta t, b = \alpha + \frac{i}{2}\omega_0\delta t$$

$$E_{1/2} = \frac{\Gamma(1/2)\Gamma(b-a)}{\Gamma(b)\Gamma(1/2-a)}, \tilde{E}_{3/2} = \frac{\Gamma(3/2)\Gamma(b-a)}{\Gamma(b+1/2)\Gamma(1-a)}$$

$$E'_c = E_c(a \leftrightarrow b)$$



- The second is a End-Critical protocol (ECP)

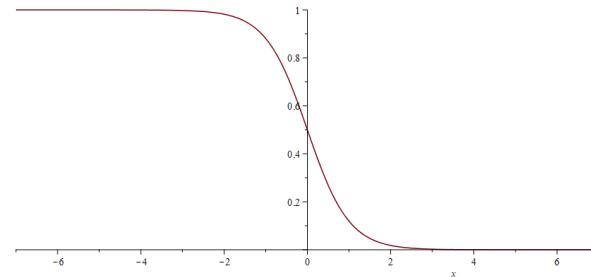
$$\omega^2(\tau) = \frac{1}{2}\omega_0^2 \left(1 - \tanh \frac{\tau}{\delta t}\right)$$

- The “in” classical solution is

$$f_{ECP} = \frac{1}{\sqrt{2\omega_{in}}} \exp[-i\omega_+ \tau - i\omega_- \delta t \log(2\cosh(\tau/\delta t))]$$

$${}_2F_1[1 + i\omega_- \delta t, i\omega_- \delta t; 1 - i\omega_{in} \delta t; \frac{1}{2}(1 + \tanh(\tau/\delta t))]$$

$$\omega_{\pm} = \frac{1}{2}(\omega_{out} \pm \omega_{in})$$



Scaling for CCP

- At times in the middle of the quench or just when the quench ends we find different scaling regimes.
- We find answers consistent with Kibble Zurek scaling for $\omega_0\delta t \gg 1$

$$\langle O(0) \rangle \sim \frac{\sqrt{\pi}}{2} N^2 \sqrt{\frac{\delta t}{\omega_0}}$$

- In the fast regime $\omega_0 t \ll \omega_0 \delta t \ll 1$

$$\langle \mathcal{O}(t) \rangle \approx \frac{N^2}{2\omega_0} \left\{ 1 + 2\log 2 \cdot \omega_0^2 \delta t^2 + 2\omega_0^2 \delta t \cdot t + \mathcal{O}(\omega_0^4 \delta t^4, \frac{t^2}{\delta t^2}) \right\}$$

- This is consistent with perturbation theory. However unlike relativistic theories the linear response term is not proportional to ω_0^2

Early time response : ECP

- For end critical protocols it is more appropriate to look at response for a range of times. Now at any given time a sufficiently slow quench will always be adiabatic – since the gapless point is approached asymptotically.
- For slightly higher quench rates there is a Kibble-Zurek regime where

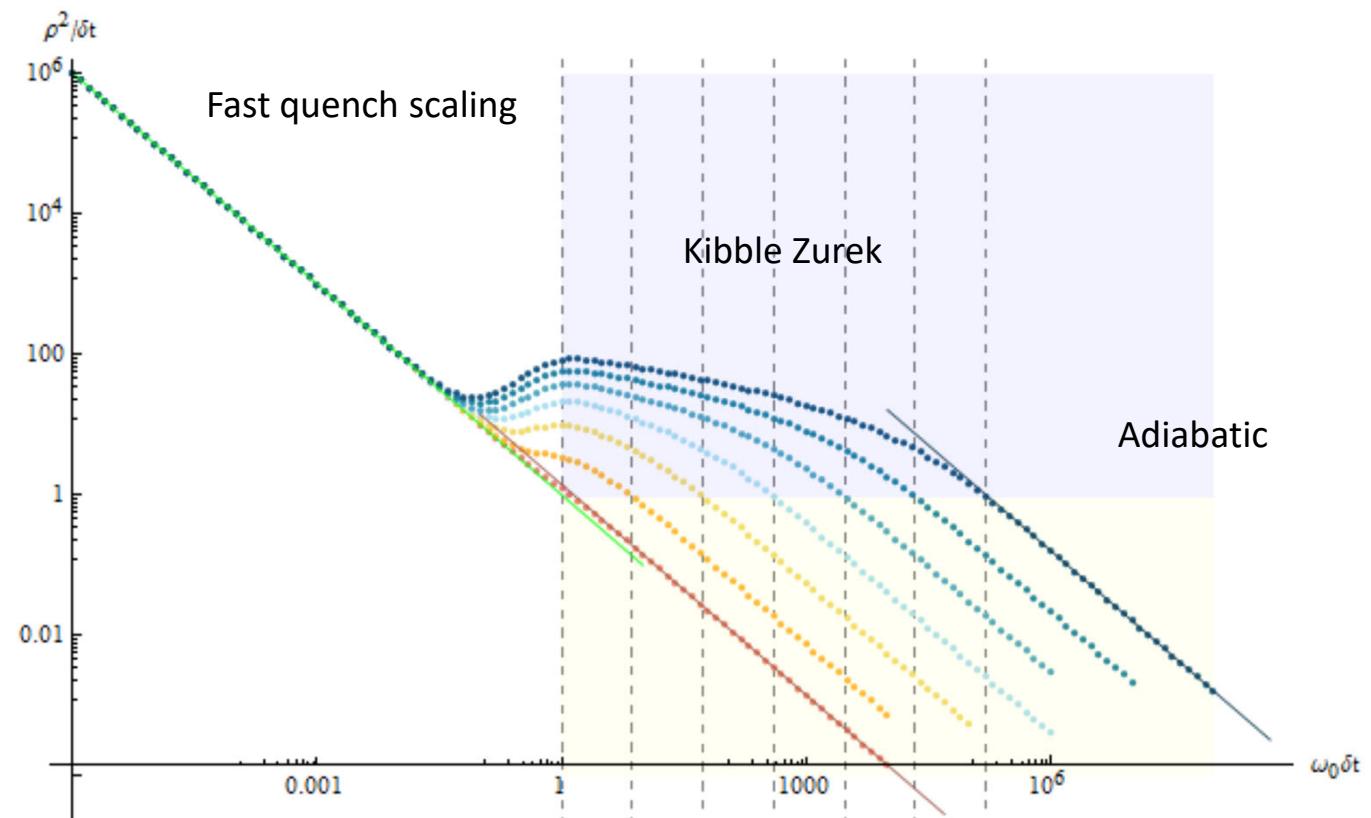
$$\langle \mathcal{O} \rangle \sim N^2 \rho(t_{KZ})^2 = \frac{N^2}{\omega_{KZ}} \sim N^2 \delta t$$

- An expansion of the exact answer actually reveals a log correction

$$\rho^2(\tau) \sim \delta t \left[\frac{2}{\pi} \left(-\log \omega_0 \delta t + \log 2 - \gamma_E + \frac{\tau}{\delta t} \right)^2 + \frac{\pi}{2} \right] \sim \mathcal{O}(1)$$

- In the fast quench regime $\omega_0\delta t \ll \omega_0\tau \ll 1$. we get

$$\rho^2(\tau) \sim \omega_0\delta t^2 \left(-\frac{\zeta(3)}{4}\omega_0^2\delta t^2 + \frac{\tau}{\delta t} \right)^2 + \frac{1}{\omega_0} = \frac{1}{\omega_0} + \omega_0\tau^2 - \frac{\zeta(3)}{2}(\omega_0\delta t)^3\tau.$$



Connection to Two Dimensional Strings

- Fermions in an **inverted harmonic oscillator** potential is well known to provide a non-perturbative description of **2 dimensional string theory** – via its connection to **Matrix Quantum Mechanics**.
- In the double scaled limit the Hamiltonian is given by

$$H = \frac{1}{2} \int dx \psi^\dagger \left[-\frac{1}{g_s} \partial_x^2 - g_s x^2 \right] \psi$$

- Where g_s is the scaled **string coupling** which is held fixed in the large N double scaled limit. (The physical string coupling is position dependent, but proportional to g_s)
- A **time-dependent coupling** is then the problem with a **time dependent mass** and a **constant frequency**.
- In this case we get interesting examples of two dimensional **cosmologies**.

Epilouge

- Our results show that there is interesting universal behavior for quench rates at **intermediate scales** between the **UV** and **physical mass scales**.
- This result arose from an interesting interplay of holography and standard field theoretic techniques.
- We have mostly concentrated on the behavior near the ***end-point*** of the quench. While this gives late time results for quantities like the energy density- it would be interesting to explore how this universality manifests in other quantities at **late times**, and their implications to **thermalization**.
- In particular, it would be interesting to see how this appears in the **time evolution of entanglement** which has been so far studied for abrupt changes.

Holographic Cosmologies

- In the holographic context, there are situations where a quantum quench on the boundary does not lead to black hole formation.
- Rather, the bulk space-time develops a space-like region of high curvature akin to **cosmological singularities**. ([A. Awad, S.R.D., K. Narayan, S. Trivedi : 2008,2009](#)).
- One example is the **AdS-Kasner**

$$ds_b^2 = \tilde{g}_{\alpha\beta} dx^\alpha dx^\beta = -dt^2 + \sum_{a=1}^{(d-1)} t^{2p_a} (dx^a)^2,$$

$$\Phi(t) = \alpha \log |t|$$

$$\sum_{a=1}^{d-1} p_a = 1, \quad \sum_{a=1}^{d-1} p_a^2 = 1 - \frac{\alpha^2}{2}$$

- A time dependent dilaton means that the coupling of the boundary theory is time dependent – so a **quantum quench**. There are also solutions where the dilaton is bounded

$$ds^2 = \frac{1}{z^2} dz^2 + \left[|\sinh(2t)| \left(-dt^2 + \frac{dr^2}{1+r^2} + r^2 d\Omega_2^2 \right) \right]$$

$$\Phi = -2\sqrt{3}\tanh^{-1}(e^t).$$

- These backgrounds were studied in the past with the hope that the boundary theory provides a perfectly well-defined time evolution, thus “resolving” the singularity.
- **There was no definitive conclusion about this.**
- In recent years this problem has been revisited in several contexts. The aim has been to find **what are the signatures of such a bulk singularity in the boundary theory**.
- ***Englehardt & Horowitz (2014,2015,2016); Barbon & Rabinovici (2015,2018); Frenkel, Hartnoll, Kruthoff & Shi (2020); Hartnoll, Horowitz, Kruthoff & Santos (2021).***

- One interesting question concerns the behavior of complexity of the quantum state near the singularity.
- We recently investigated this issue using a particular notion of complexity.
- The main result is that the complexity decreases as one approaches the singularity in a universal manner – independent of the Kasner exponents. (P. Caputa, D. Das and S.R.D. (2022)).
- This is consistent with the fact that the dimensionality of the Hilbert space decreases as one approaches the singularity.

Thank You