# **Basics of** *v***eutrino Oscillations**

#### *Mustafa Ashry* mustafa@sci.cu.edu.eg

Department of Mathematics, Faculty of Science, Cairo University



BUE Summer School and Internship 2022



21 July 2022

## **1** $\nu$ eutrino Identity

- **2**  $\nu$ eutrino Masses and Mixing
- **B** *v*eutrino Oscillations
- 4  $\nu$ eutrino Open Questions
- **5** *v*eutrino Seesaw Mechanism

- **1**  $\nu$ **eutrino Identity**
- **2**  $\nu$ eutrino Masses and Mixing
- **B**  $\nu$ eutrino Oscillations
- **4** *ν*eutrino Open Questions
- **5** *v*eutrino Seesaw Mechanism

- **1**  $\nu$ **eutrino Identity**
- **2**  $\nu$ eutrino Masses and Mixing
- **3**  $\nu$ **eutrino Oscillations**
- **4** *v*eutrino Open Questions
- **5** *v*eutrino Seesaw Mechanism

- **1**  $\nu$ **eutrino Identity**
- **2**  $\nu$ eutrino Masses and Mixing
- **3** veutrino Oscillations
- **4** *v***eutrino Open Questions**
- **5** *v*eutrino Seesaw Mechanism

- **1**  $\nu$ **eutrino Identity**
- **2**  $\nu$ eutrino Masses and Mixing
- **3** *v***eutrino Oscillations**
- **4** *v***eutrino Open Questions**
- **5** *v***eutrino Seesaw Mechanism**

 $-\nu$ eutrino Identity

## Outline

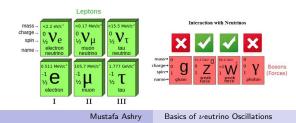
## **1** $\nu$ **eutrino Identity**

- **2** veutrino Masses and Mixing
- **3** veutrino Oscillations
- **4**  $\nu$ eutrino Open Questions
- **5** *v***eutrino Seesaw Mechanism**

## Identity

- $\nu$ eutrinos are neutral spin- $\frac{1}{2}$  elementary fermions.
- veutrinos are leptons: interact in the Standard Model (SM) via weak interactions only.
- $\nu$ eutrinos exist in three flavors: electron- $\nu$ eutrino, muon- $\nu$ eutrino and tau- $\nu$ eutrino:  $\nu_{\ell} = (\nu_e, \nu_{\mu}, \nu_{\tau})$ .
- $\nu$ eutrino flavors are created in weak interactions in association with the corresponding charged lepton  $\ell = (e, \mu, \tau)$ .

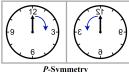
•  $\nu$ eutrinos are tiny massive  $\mathcal{O}(eV)$ .

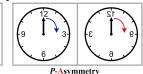


Basics of $\nu$ eutrino Oscillations	
$-\nu$ eutrino Identity	

## Parity

- Parity is the transformation under space reflection.
- Parity was assumed at the beginning to be a symmetry of nature (1927, Eugene Wigner).
- That's a 'mirrored' image of a natural system behaves in the same way as the 'mirror' image of that system does.



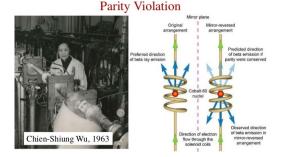


Fermions are chiral. They are absolutely classified into left- and right-handed chiralities. They are helicial too, according to their relative spin projection in the direction of motion.



## **Parity Violation**

- Parity is a symmetry for both strong and electromagnetic interactions.
- In 1963, Wu found that Parity is maximally violated in the weak interactions.
- Only left-handed fermions feel the weak interactions.



• (Known)  $\nu$  eutrinos are only left-handed.

-veutrino Masses and Mixing

## Outline

7/28

## **1** $\nu$ eutrino Identity

- **2**  $\nu$ eutrino Masses and Mixing
- **3** *v*eutrino Oscillations
- **4** *v*eutrino Open Questions
- **5** *v*eutrino Seesaw Mechanism

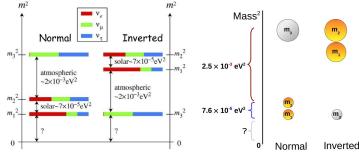
veutrino Masses and Mixing

#### veutrino mass hierarchies

veutrinos masses are of order of little eV's.



Limits from solar and atmoshperic νeutrino experiments propose the normal and inverted hierarchies for the νeutrino masses.



#### $\nu$ eutrino mixing

- veutrinos nonzero masses yields to veutrino mixing phenomena [1].
- That is, veutrino flavors  $\nu_e$ ,  $\nu_\mu$ ,  $\nu_\tau$  do not have definite masses.
- Instead, νeutrino flavors are linear combinations of three other mass states ν<sub>1</sub>, ν<sub>2</sub>, ν<sub>3</sub> with definite masses m<sub>1</sub>, m<sub>2</sub>, m<sub>3</sub>.

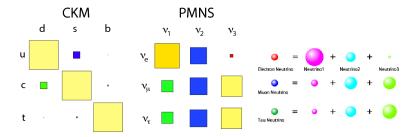
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{bmatrix} = \begin{bmatrix} 0.82 \pm 0.01 & 0.54 \pm 0.02 & -0.15 \pm 0.03 \\ -0.35 \pm 0.06 & 0.70 \pm 0.06 & 0.62 \pm 0.06 \\ 0.44 \pm 0.06 & -0.45 \pm 0.06 & 0.77 \pm 0.06 \end{bmatrix}$$

└─ veutrino Masses and Mixing

## **CKM & PMNS Mixing matrices**

- The CKM quark mixing matrix is almost diagonal.
- The PMNS mixing matrix of veutrinos is equilibrated; all elements have approximately the same order.



 $-\nu$ eutrino Oscillations

## Outline

**1**  $\nu$ eutrino Identity

- **2**  $\nu$ eutrino Masses and Mixing
- **3** veutrino Oscillations
- **4** *v*eutrino Open Questions
- **5** *v***eutrino Seesaw Mechanism**

#### **Beats Lagrangian**

Lagrangian for forced oscillations ( $\omega$ : free oscillation frequency) [2]

$$L = \frac{1}{2}m\dot{x}^2 - \frac{1}{2}m\omega^2 x^2 + xF(t),$$
(1)

• *Euler-Lagrange* equation

$$\ddot{x} + \omega^2 x = F(t)/m.$$
 (2)

Free oscillations (A: amplitude and  $\alpha$ : initial phase)

$$x(t) = A\cos(\omega t + \alpha), \tag{3}$$

Oscillatory force

$$F(t) = f \cos(\gamma t + \beta).$$
(4)

#### **Beats amplitude**

Near the resonance  $(\gamma = \omega)$ ,  $\gamma = \omega + \epsilon$ ,  $\frac{\epsilon}{\omega} \ll 1$ , the motion is small oscillations with variable amplitude

$$x(t) = C(t) \exp(i\omega t), \tag{5}$$

where

$$C^{2} = a^{2} + b^{2} + 2ab\cos(\epsilon t + \beta - \alpha),$$
 (6)

and a, b are constants.

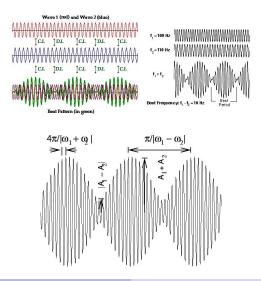
• The amplitude varies slowly with frequency  $\epsilon$  between the limits

$$|a-b| \le C \le a+b. \tag{7}$$

This phenomena is called the "beats".  $\epsilon$  is the beat frequency.

 $-\nu$ eutrino Oscillations

Click BEATS Enjoy!



Mustafa Ashry Basics of veutrino Oscillations

#### $\nu$ eutrino two flavor mixing

Mixing only two flavors  $u_lpha,
u_eta$  of two mass states  $u_i,
u_j$ 

$$\nu_{\alpha} = \nu_i \cos \theta_{ij} + \nu_j \sin \theta_{ij}, \tag{8}$$

$$\nu_{\beta} = -\nu_i \sin \theta_{ij} + \nu_j \cos \theta_{ij}, \tag{9}$$

- When a  $|\nu_{\alpha}(0), \vec{p}\rangle$  veutrino is produced with momentum  $\vec{p}$  at time t = 0, the  $\nu_i$  and  $\nu_j$  components will have slightly different energies  $E_i$  and  $E_j$  due to their slightly different masses.
- In QM, their associated waves have slightly different frequencies, and their interference gives rise to the 'beats' phenomenon.
- As a result, the original beam of ν<sub>α</sub> develops a ν<sub>β</sub> component whose intensity oscillates as it travels through space, meanwhile, the intensity of the ν<sub>α</sub> νeutrino beam itself is correspondingly reduced.
- This is the 'veutrino oscillations' phenomenon and its occurence follows from simple QM.

#### Propagation

At time t = 0, the initial state (8) is produced with momentum  $\vec{p}$ 

$$|\nu_{\alpha}(0),\vec{p}\rangle = |\nu_{i},\vec{p}\rangle\cos\theta_{ij} + |\nu_{j},\vec{p}\rangle\sin\theta_{ij}.$$
 (10)

After time t, Schrödinger equation for the mass states  $\nu_i, \nu_j$  dictates

$$\nu_{\alpha}(t), \vec{p} \rangle = a_i(t) |\nu_i, \vec{p}\rangle \cos\theta_{ij} + a_j(t) |\nu_j, \vec{p}\rangle \sin\theta_{ij}, \qquad (11)$$

where

$$a_{i,j}(t) = \exp\left(-iE_{i,j}t\right) \tag{12}$$

are the usual oscillating time factors associated with any quantum mechanical stationary state.

### Propagation

Inverting eqs. (8,9)

$$\nu_i = \nu_\alpha \cos \theta_{ij} - \nu_\beta \sin \theta_{ij},\tag{13}$$

$$\nu_j = \nu_\alpha \sin \theta_{ij} + \nu_\beta \cos \theta_{ij}.$$
 (14)

Substituting into eq. (11) gives the oscillation formula

$$|\nu_{\alpha}(t),\vec{p}\rangle = A(t)|\nu_{\alpha}(0),\vec{p}\rangle + B(t)|\nu_{\beta}(0),\vec{p}\rangle,$$
(15)

where

$$A(t) = a_i(t)\cos^2\theta_{ij} + a_j(t)\sin^2\theta_{ij},$$
(16)

$$B(t) = \sin \theta_{ij} \cos \theta_{ij} [a_j(t) - a_i(t)].$$
(17)

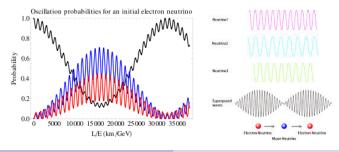
### **Oscillation Probability**

• The probability of finding a  $\nu_{\beta}$  state is therefore

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = |\langle \nu_{\beta}(0), \vec{p} | \nu_{\alpha}(t), \vec{p} \rangle|^{2}$$
$$= |B(t)|^{2} = \sin^{2}(2\theta_{ij}) \sin^{2}[\frac{1}{2}(E_{j} - E_{i})t].$$
(18)

• The probability of finding a  $\nu_{\alpha}$  state is therefore

$$P(\nu_{\alpha} \to \nu_{\alpha}) = |\langle \nu_{\alpha}(0), \vec{p} | \nu_{\alpha}(t), \vec{p} \rangle = |A(t)|^2 = 1 - |B(t)|^2.$$
(19)



Basics of  $\nu$ eutrino Oscillations

## Remarks

- No oscillation with vanishing mixing  $\theta_{ij} = 0$ .
- For small mixing  $\theta_{ij}$  and large energy difference, oscillation is negligible.
- These formulas assume that the veutrinos are propagating in a vacuum. This is usually a very good approximation, because of the enormous mean free paths for veutrinos to interact with matter.
- Mikheyev-Smirnov-Wolfenstein (MSW) effect shows that veutrino oscillations can be enhanced when veutrinos traverse very long distances in matter due to weak interactions with matter's electrons analogous to the electromagnetic process leading to the refractive index of light in a medium.
- The MSW effect was dramatically confirmed in the 'Sudbury veutrino Observatory (SNO)', and resolved the solar veutrino problem.

Basics of $\nu$ eutrino Oscillations
└ veutrino Oscillations

#### Detection

- The time t traveled by a veutrino is determined by the distance L of the veutrino detector its source.
- veutrino are always ultra-relativistic and their momenta are much greater than their possible masses and they approximatly travel at the speed of light.
- So  $t \approx L$  and

$$E_j - E_i = (m_j^2 + p^2)^{1/2} - (m_i^2 + p^2)^{1/2} \approx \frac{m_j^2 - m_i^2}{2p}.$$
 (20)

Thus the oscillation probability (18) can be written as

$$P(\nu_{\alpha} \to \nu_{\beta}) = \sin^2(2\theta_{ij}) \sin^2[L/L_0], \qquad (21)$$

where the oscillation length

$$L_0 = \frac{4E}{m_j^2 - m_i^2}.$$
 (22)

## Celebration

- The oscillation lengths are typically of order 100 km or more.
- Oscillations can be safely neglected under normal laboratory conditions.
- The Nobel Prize in Physics 2015 for the discovery of *v*eutrino oscillations. Click ▶ NPNO.pdf.



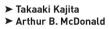


Photo © Takaaki Kajit Takaaki Kajita Prize share: 1/2



Photo: K. MacFarlane. Queen's University /SNOLAB Arthur B. McDonald Prize share: 1/2

## The Nobel Prize in Physics 2015



"for the discovery of neutrino oscillations" -veutrino Open Questions

## Outline

**1**  $\nu$ eutrino Identity

- **2**  $\nu$ eutrino Masses and Mixing
- **3** *v*eutrino Oscillations
- **4**  $\nu$ **eutrino Open Questions**
- **5** *v*eutrino Seesaw Mechanism

## **Open Problems**

- veutrino mass heirarchy. Why veutrino masses are tiny?
- veutrino absolute masses. What they are?
- veutrino: Majorana? Are veutrinos their own antiparticles?



└─ νeutrino Seesaw Mechanism

## Outline

**1**  $\nu$ eutrino Identity

- **2**  $\nu$ eutrino Masses and Mixing
- **B** *v*eutrino Oscillations
- **4**  $\nu$ eutrino Open Questions

## **5** *v***eutrino Seesaw Mechanism**

## Type I seesaw

It is thought that veutrinos not only mix among each others, but also with their right-handed much heavy cousins which balance their tiny masses on seesaw.



■ SM+type I seesaw Lagrangian

$$\mathcal{L}_{\nu} = Y_{ij}^{\nu} \bar{\nu}_{L_i} (v+h) \nu_{R_j} + M_{R_{ij}} \bar{\nu}_{R_i}^c \nu_{R_j} + h.c.$$
(23)

• Light and heavy  $\nu$ eutrino masses for  $M_R \gg$  Higgs vev (v)

$$m_{\nu_{\ell}} = \frac{m_D^2}{M_R}, \quad m_{\nu_h} \approx M_R, \quad m_D = Y^{\nu} v$$
(24)

### Assignment to think

- What happens to the veutrino oscillation pattern at very large time scale compared to the oscillation scale itself?
- Quarks mix too via the CKM matrix. Are there quark oscillations?
- Assume there are quark oscillations. On which scale they would be?
- Read about *Wu* experiment.



νeutrino Seesaw Mechanism

Thank you! Questions?

└─ *v*eutrino Seesaw Mechanism

### References

- [1] B. Martin and G. Shaw, <u>Particle physics</u>. John Wiley & Sons, 2013.
- [2] L. D. Landau and E. M. Lifshits, <u>Quantum Mechanics</u>, vol. v.3 of <u>Course of Theoretical Physics</u>. Butterworth-Heinemann, Oxford, 1991.