

Misner string:

TN metric has a string-like singularity along z-axis to see that let us consider a family of time-like surface generated by surface $f = t$, which gives $\nabla_\mu t = \delta_{\mu\nu}$. Its norm is

$$|\mathcal{J}|^2 = |\nabla t|^2 = \frac{4n^2}{(r^2 - n^2)} \frac{\cos^2 \theta}{\sin^2 \theta} + \frac{1}{f(r)}.$$

* This shows two singular behaviors, one at $f(r) = 0$, which is a coordinate singularity and at $\theta = 0, \pi$ which cannot be removed at the same time using proper coordinates! Metric is

$$dS^2 = -f(dt - 2n(\cos\theta + 1)d\varphi)^2 + f^{-1}dr^2 + \dots$$

set $c = -1$.

$$dS^2 = -f(dt - 2n(\cos\theta - 1)d\varphi)^2 + \dots$$

set $c = +1$.

$$dS^2 = -f(dt - 2n(\cos\theta + 1)d\varphi)^2 + \dots$$

* Two patches can cover manifold smoothly only if

$$\Rightarrow \Delta t^{(\pm)} = 4n \overbrace{\Delta \varphi}^{2\pi} \Rightarrow \boxed{\beta = 8\pi n}.$$

$$\star \text{But for our soln. } \beta = 4\pi r_0 = 8\pi n \Rightarrow \boxed{r_0 = 2n} \\ \boxed{M = \frac{3}{4}n}.$$

* This bring restriction to thermodynamics which affect the existence of independent terms like $\oint d\mu$ which could contribute to BH thermodynamics?

TN Unconstrained Thermodynamics

* The price for removing Misner string is Constraining Thermodynamics

$$\beta(r_0) = 8\pi n.$$

* As a result entropy $S \propto A_h$ (not horizon area)

• mass has to be $M = \frac{3}{4}n$

• $T \xrightarrow{n \rightarrow 0} \neq T_{\text{sch}}$ (it diverges in this limit!)

* This situation lead a group of authors to relax the above cond.

* Their idea was to add a pair of thermody. quantity (X, Y) such that

$$dU = TdS + \phi_e d\phi_e + \underline{X dY} \quad (\text{two possible interpretations})$$

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$$X dY = T_{MS} dS_{MS}$$

$$\text{but this} \Rightarrow \beta_{MS} = \beta = 8\pi n \\ (\text{back to } \square 1.)$$

②

$$X dY = \phi_N dN \quad \text{where } N \text{ is a conserved quant.}$$

* We proposed the second choice with $N = n$ and $\phi_N = -\frac{n}{2r_s}$. This is true in TN-Minkowski

* for TN-AdS $N = n(1 + \frac{4n^2}{c^2})$, $\Lambda = -3/c^2$, (c AdS radius)

* In the coming pages we will talk more about TN-dyons in Minkowski space.

Conserved charges for TN-Minko

* To calculate total mass we use Komar integral with $\chi^{(t)}$,

$$\left(-\frac{1}{4\pi}\right) (\star d\chi^{(t)}) = \frac{1}{4\pi} \left(\frac{4n^2 f}{r^2+n^2} \cos\theta dr \wedge d\phi + f(r^2+n^2) \sin\theta d\theta \wedge d\phi \right)$$

Masses: $M^\infty = M$ $S M^{(+)} = M^{(+)} = -n\phi_n$,
 $M^h = M - 2n\phi_n$

NUT charges:

$$N^\infty = n \quad N^h = 0$$

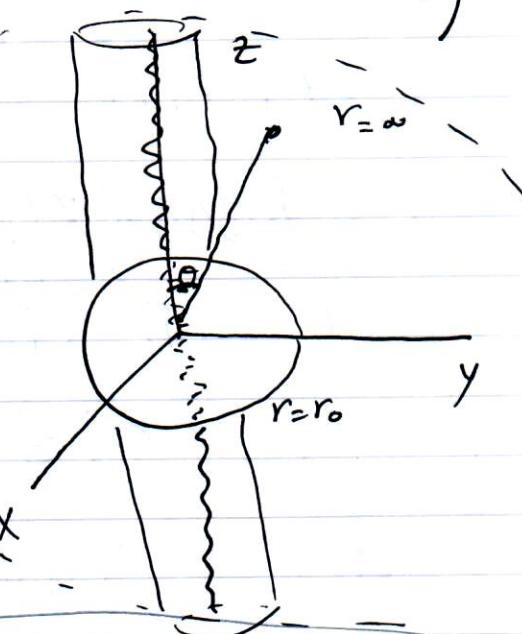
$$N^+ = N^- = \frac{n}{2}$$

E.M. charges: $Q_e^\infty = q_e$ $S Q_e^h = q - 2n\phi_m$

$$Q_m^\infty = q_m \quad S Q_m^h = q_m + 2n\phi_e$$

There exist

* Important: ~~not~~ charges along z -axis and they are indep of having Misner string or not!



* 1st Law $d(M - n\phi_n) = TdS + \phi_n dn + \phi_e dQ_e^\infty + \phi_m dQ_m^h$.

* internal energy is not the mass, but $U = M - n\phi_n > 0$, which is telling us that there is another source for the total energy which is the nut charge besides the mass "M"! Helmholtz free energy

* On-shell gravit. action $I_g = I(T, n) = \beta \bar{F}$

$$\bar{F} = \frac{I}{\beta} = U - TS.$$

* To understand more about BH Thermodynamics and their ensembles we need to discuss Euclidean Path Integral since it is the formal Theory for calculating those quantities.

In Q.M. Path integral

$$\langle q_2(t_2) | q_1(t_1) \rangle = \int D[q] e^{i I(q)}$$


In Q.F.T. $I(\phi)$

$$= \langle \phi_2, t_2 | \phi_1, t_1 \rangle = \int D[\phi] e^{-I(\phi)}$$

$$t_2 = t, t_1 = 0, g = \gamma$$

$$Z = \text{Kern}(\phi_2, t; \phi_1, 0) = \langle \phi_2 | e^{-Ht} | \phi_1 \rangle$$

$$t = -\beta \text{ (Eucl.)} \Rightarrow \text{Kern}(\phi_2, \phi_1) = \delta(\phi_2 - \phi_1)$$

$$\sum_{\phi} \langle \phi | e^{-\beta H} | \phi \rangle = \text{tr } e^{-\beta H} = Z_{\text{part.}}$$

$$Z = \int D[g, \phi] e^{-I^{\text{ext}}(g, \phi)} = \text{tr } e^{-\beta H(g, \phi)}$$

In gravity β is not arb. since regularity of the fields at boundary requires certain temp.!

$$ds^2 = -f dt^2 + f^{-1} dr^2 + r^2 d\Omega^2, \quad f = 1 - \frac{2M}{r}$$

$$f = f(r) + (r - r_s) f' + \dots$$

$$ds^2 = (r - r_s) f' d\bar{t}^2 + \frac{dr^2}{(r - r_s) f'} + \dots$$

$$\int dR = \sqrt{\frac{dr}{(r - r_s) f'}} \Rightarrow R = \sqrt{\frac{2M}{f'} \sqrt{r - r_s}} = \frac{2\sqrt{(r - r_s)}}{\sqrt{f'}} \Rightarrow R^2 f' = (r - r_s)$$

$$ds^2 = \frac{R^2}{4} f'^2 d\bar{t}^2 + \frac{dR^2}{R^2 f'^2} + R^2 d\phi^2 \quad \Delta \bar{t} = \beta = T^{-1}$$

$$R^2 d(f' \bar{t}) \Rightarrow \bar{t} = \frac{1}{2} \Delta \bar{t} = \frac{1}{2} \frac{4\bar{n}}{f'} \Rightarrow \Delta \bar{t} = \frac{4\bar{n}}{f'}$$

Thermodynamic Ensembles & Regularity Conditions

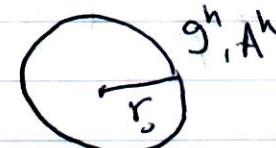
BH's Thermodynamics is defined formally through Euclidean Path Integrals, or

$$Z = \int D[g, A] e^{-I(g, A)}$$

where g and A are metric and matter field components.

* As a boundary condition values of g, A must be fixed at the boundary.

* We consider all possible values between horizon and $r \rightarrow \infty$. But g, A should be finite at the two boundaries!



* As we take $t \rightarrow i\tau$, we get a conical singularity at the horizon unless $[t] = \beta = \frac{4\pi}{f'(r_0)}$.

* Norm of A , or $|A|^2 = A_\mu A^\mu$, must be finite at the horizon, e.g.

for R.N. soln $|A|^2 \sim -\frac{A_0^2}{f(r)} = \left(-\frac{q}{r_0} + c\right)^2 \cdot \frac{1}{f(r)}$ (at horizon!)

$\Rightarrow C = \frac{q}{r_0} = \phi_e$ (electric Pot.) $\xrightarrow{\text{BH f.t.}} \text{BH f.t.}$

At $r=\infty$ $A_0 \rightarrow \phi_e$ (bdry cond.) $\Rightarrow dM = TdS + \phi_e dQ_e$

$Z = \begin{cases} \text{Tr } e^{-\beta H} \simeq e^{-I_{\text{on-shell}}} = e^{-\beta F} \xrightarrow{\text{Helm. free. Eng.}} \\ \text{Tr } e^{-\beta(H - \mu_i C_i)} \simeq e^{-I_{\text{on-shell}}} = e^{-\beta G} \xrightarrow{\text{Gibbs free. Eng.}} \end{cases}$

Work term $\phi_e dQ_e = e \phi_e dN = \mu dN$

Sch. soln $I(T) = \beta F$ (Cannonical ens.)

R.N. soln $I(T, \phi_e) = \beta G$ (grand-can ens.)

TN $I(T, n) = \beta F$ (Can. ens.)

$$A_{\mu}^{ek} = \left(-\frac{q}{r} + C, 0, 0, 0 \right) \xrightarrow[r \rightarrow \infty]{r \rightarrow r_0} \left(C, 0, 0, 0 \right)$$

fixed
+ finite

$$\xrightarrow[r \rightarrow r_0]{} \left(-\frac{q}{r_0} + C, 0, 0, 0 \right)$$

fixed
+ finite

$$A_{\mu}^{\text{mag-ch.}} = \left(0, 0, 0, q_m (\cos \theta + c') \right) \xrightarrow[r \rightarrow \infty]{} A_{\mu}^{(m)\infty} = \left(0, 0, 0, q_m (\cos \theta) \right)$$

must be fixed

$\xrightarrow[r \rightarrow r_0]{} \rightarrow$ already regular.

$$\xrightarrow[\text{dir.}]{} |\vec{A}| = q_m \frac{(\cos \theta + c')^2}{r^2 \sin^2 \theta}$$

$$\Rightarrow A_q^{(\pm)} = q_m (\cos \theta \pm 1)$$

TN Thermodynamics (our approach):

- * Relaxing the cond. $\beta = 8\pi n$ and dealing with "n" as a conserved charge as well as Considering $V = M - n\phi_n$ is

The essence of our approach to construct a consistent Thermodynamics of TN.

$$\star V = -\partial_\beta \ln Z = \partial_\beta I_{\text{on-shell}} = M - n\phi_n.$$

$$dV = TdS + \phi_n dn + \phi_e dQ_e^\infty + \phi_m dQ_m^h.$$

- * But the first law is not enough as a consistency check to our approach; but we have also two other important relations

* Gibbs-Duhem relation

$$\frac{\partial I_{\text{on-shell}}}{\partial \beta} = G = M - n\phi_n - TS - Q_e\phi_e$$

* Smarr's Relation

$$U = M - n\phi_n = TS + n\phi_e + Q_e\phi_e + Q_m\phi_m$$

(non-trivial !)

* 1st Law

$$dU = TdS + \phi_n dn + \phi_e dQ_e^\infty + \phi_m dQ_m^h.$$